

```
In[ ]:= Get@FileNameJoin@{NotebookDirectory[], "BlackHoleEquations.m"};
```

# Stefan's modified black hole equations

## Connection to Part III project

The purpose of this script is to give a concrete set of tools for extracting the reduced field equations in a static, spherical spacetime. With these tools, it is hoped that the remainder of the project time can be spent on analysis of the physics.

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Package xAct`xPerm` version 1.2.3, {2015, 8, 23}

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Connecting to external linux executable...

Connection established.

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Package xAct`xTensor` version 1.2.0, {2021, 10, 17}

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Package xAct`xPerm` version 1.2.3, {2015, 8, 23}

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 Package xAct`xPert` version 1.0.6, {2018, 2, 28}

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 and Guillermo A. Mena Marugan, under the General Public License.

\*\* Variable \$PrePrint assigned value ScreenDollarIndices

\*\* Variable \$CovDFormat changed from Prefix to Postfix

\*\* Option AllowUpperDerivatives of ContractMetric changed from False to True

\*\* Option MetricOn of MakeRule changed from None to All

\*\* Option ContractMetrics of MakeRule changed from False to True

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 Package xAct`Invar` version 2.0.5, {2013, 7, 1}

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 D. Yllanes and R. Portugal, under the General Public License.

\*\* DefConstantSymbol: Defining constant symbol sigma.

\*\* DefConstantSymbol: Defining constant symbol dim.

\*\* Option CurvatureRelations of DefCovD changed from True to False

\*\* Variable \$CommuteCovDsOnScalars changed from True to False

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 Package xAct`xCoba` version 0.8.6, {2021, 2, 28}

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 Jose M. Martin-Garcia, under the General Public License.

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 Package xAct`SymManipulator` version 0.9.5, {2021, 9, 14}

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 Package xAct`xTras` version 1.4.2, {2014, 10, 30}

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\*\* Variable \$CovDFormat changed from Postfix to Prefix

\*\* Option CurvatureRelations of DefCovD changed from False to True

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 -----

\*\* DefManifold: Defining manifold M4.

\*\* DefVBundle: Defining vbundle TangentM4.

\*\* DefTensor: Defining symmetric metric tensor G[-a, -b].

\*\* DefTensor: Defining antisymmetric tensor epsilonG[-a, -a1, -b, -b1].

\*\* DefTensor: Defining tetrametric TetraG[-a, -a1, -b, -b1].

\*\* DefTensor: Defining tetrametric TetraG†[-a, -a1, -b, -b1].

\*\* DefCovD: Defining covariant derivative CD[-a].

\*\* DefTensor: Defining vanishing torsion tensor TorsionCD[a, -a1, -b].

\*\* DefTensor: Defining symmetric Christoffel tensor ChristoffelCD[a, -a1, -b].

\*\* DefTensor: Defining Riemann tensor RiemannCD[-a, -a1, -b, -b1].

\*\* DefTensor: Defining symmetric Ricci tensor RicciCD[-a, -a1].

\*\* DefCovD: Contractions of Riemann automatically replaced by Ricci.

\*\* DefTensor: Defining Ricci scalar RicciScalarCD[].

\*\* DefCovD: Contractions of Ricci automatically replaced by RicciScalar.

\*\* DefTensor: Defining symmetric Einstein tensor EinsteinCD[-a, -a1].

\*\* DefTensor: Defining Weyl tensor WeylCD[-a, -a1, -b, -b1].

\*\* DefTensor: Defining symmetric TFRicci tensor TFRicciCD[-a, -a1].

\*\* DefTensor: Defining Kretschmann scalar KretschmannCD[].

\*\* DefCovD: Computing RiemannToWeylRules for dim 4

\*\* DefCovD: Computing RicciToTFRicci for dim 4

\*\* DefCovD: Computing RicciToEinsteinRules for dim 4

\*\* DefTensor: Defining symmetrized Riemann tensor SymRiemannCD[-a, -a1, -b, -b1].

\*\* DefTensor: Defining symmetric Schouten tensor SchoutenCD[-a, -a1].

\*\* DefTensor: Defining symmetric cosmological Schouten tensor SchoutenCCCD[LI[], -a, -a1].

```

** DefTensor: Defining symmetric cosmological Einstein tensor EinsteinCCCD[LI[_], -a, -a1].
** DefCovD: Defining covariant derivative CD[-a]. to be symmetrizable
** DefTensor: Defining weight +2 density DetG[]. Determinant.
** DefParameter: Defining parameter PerturbationParameterG.
** DefTensor: Defining tensor PerturbationG[LI[order], -a, -a1].

```

Define a Planck mass.

```

** DefConstantSymbol: Defining constant symbol MPl.

```

Define the Schwarzschild functions.

```

** DefScalarFunction: Defining scalar function B1.
** DefScalarFunction: Defining scalar function B2.

```

Define the chart for spherical polar coordinates.

```

** DefChart: Defining chart SphericalPolar.
** DefTensor: Defining coordinate scalar ct[].
** DefTensor: Defining coordinate scalar cr[].
** DefTensor: Defining coordinate scalar ctheta[].
** DefTensor: Defining coordinate scalar cphi[].
** DefMapping: Defining mapping SphericalPolar.
** DefMapping: Defining inverse mapping iSphericalPolar.
** DefTensor: Defining mapping differential tensor diSphericalPolar[-a, iSphericalPolar a].
** DefTensor: Defining mapping differential tensor dSphericalPolar[-a, SphericalPolar a].
** DefBasis: Defining basis SphericalPolar. Coordinated basis.
** DefCovD: Defining parallel derivative PDSphericalPolar[-a].
** DefTensor: Defining vanishing torsion tensor TorsionPDSphericalPolar[a, -a1, -b].
** DefTensor: Defining symmetric Christoffel tensor
  ChristoffelPDSphericalPolar[a, -a1, -b].
** DefTensor: Defining vanishing Riemann tensor RiemannPDSphericalPolar[-a, -a1, -b, b1].
** DefTensor: Defining vanishing Ricci tensor RicciPDSphericalPolar[-a, -a1].
** DefTensor: Defining antisymmetric +1 density etaUpSphericalPolar[a, a1, b, b1].
** DefTensor: Defining antisymmetric -1 density etaDownSphericalPolar[-a, -a1, -b, -b1].

```

Set the components of the metric to those of Schwarzschild.

Added independent rule  $\dot{g}_{00} \rightarrow \Theta[r]^2$  for tensor G

Added independent rule  $\dot{g}_{01} \rightarrow 0$  for tensor G

Added independent rule  $\dot{g}_{02} \rightarrow 0$  for tensor G

Added independent rule  $\dot{g}_{03} \rightarrow 0$  for tensor G

Added dependent rule  $\dot{g}_{10} \rightarrow \dot{g}_{01}$  for tensor G

Added independent rule  $\dot{g}_{11} \rightarrow -\Omega[r]^2$  for tensor G

Added independent rule  $\dot{g}_{12} \rightarrow 0$  for tensor G

Added independent rule  $\dot{g}_{13} \rightarrow 0$  for tensor G

Added dependent rule  $\dot{g}_{20} \rightarrow \dot{g}_{02}$  for tensor G

Added dependent rule  $\dot{g}_{21} \rightarrow \dot{g}_{12}$  for tensor G

Added independent rule  $\dot{g}_{22} \rightarrow -r^2$  for tensor G

Added independent rule  $\dot{g}_{23} \rightarrow 0$  for tensor G

Added dependent rule  $\dot{g}_{30} \rightarrow \dot{g}_{03}$  for tensor G

Added dependent rule  $\dot{g}_{31} \rightarrow \dot{g}_{13}$  for tensor G

Added dependent rule  $\dot{g}_{32} \rightarrow \dot{g}_{23}$  for tensor G

Added independent rule  $\dot{g}_{33} \rightarrow -r^2 \text{Sin}[\theta]^2$  for tensor G

**\*\* DefTensor:** Defining weight +2 density DetGSphericalPolar[]. Determinant.

**\*\* DefTensor:** Defining tensor ChristoffelCDPDSphericalPolar[a, -a1, -b].

Just to check the line element which we built here, let's have a look!

$$\dot{g}_{\alpha\beta}$$

(1)

$$\begin{pmatrix} \dot{g}_{00} & \dot{g}_{01} & \dot{g}_{02} & \dot{g}_{03} \\ \dot{g}_{10} & \dot{g}_{11} & \dot{g}_{12} & \dot{g}_{13} \\ \dot{g}_{20} & \dot{g}_{21} & \dot{g}_{22} & \dot{g}_{23} \\ \dot{g}_{30} & \dot{g}_{31} & \dot{g}_{32} & \dot{g}_{33} \end{pmatrix}$$

(2)

$$\begin{pmatrix} \Theta[r]^y & 0 & 0 & 0 \\ 0 & -\Omega[r]^y & 0 & 0 \\ 0 & 0 & -r^y & 0 \\ 0 & 0 & 0 & -r^y \sin[\theta]^y \end{pmatrix}$$

Now how about the inverse metric?

$$g^{\alpha\beta} \quad (4)$$

$$\begin{pmatrix} g^{00} & g^{01} & g^{02} & g^{03} \\ g^{10} & g^{11} & g^{12} & g^{13} \\ g^{20} & g^{21} & g^{22} & g^{23} \\ g^{30} & g^{31} & g^{32} & g^{33} \end{pmatrix} \quad (5)$$

$$\begin{pmatrix} \frac{x}{\Theta[r]^2} & 0 & 0 & 0 \\ 0 & -\frac{x}{\Omega[r]^2} & 0 & 0 \\ 0 & 0 & -\frac{x}{r^2} & 0 \\ 0 & 0 & 0 & -\frac{Csc[\theta]^2}{r^2} \end{pmatrix} \quad (6)$$

Define scalar functions of the radial coordinate scalar, and functions which will represent the ADM quantities.

```
** DefTensor: Defining tensor Phi[].
```

```
** DefTensor: Defining tensor Psi[].
```

```
** DefScalarFunction: Defining scalar function Psis.
```

```
** DefScalarFunction: Defining scalar function Phis.
```

### Connection to Part III project

Something I omitted in the video tutorial above was a discussion of how to set up reduced radial-function components of a tensor field which is not the metric. There were no such tensors in the torsionful effective theory I gave to Oliver, but of course in MoND we have the unit-timelike vector field. How to work with this?

We wish to define a unit-timelike one-form.

```
** DefTensor: Defining tensor A[-a].
```

$$\mathcal{A}_\alpha \quad (7)$$

$$\{\mathcal{A}_0, \mathcal{A}_1, \mathcal{A}_2, \mathcal{A}_3\} \quad (8)$$

We know that this one-form has the property of being unit-timelike, and that the Killing algebra is spherical. So, we define a couple of scalar functions.

**\*\* DefScalarFunction:** Defining scalar function A1.

**\*\* DefScalarFunction:** Defining scalar function A2.

Added independent rule  $\mathcal{A}_0 \rightarrow \Phi[r]$  for tensor A

Added independent rule  $\mathcal{A}_1 \rightarrow \sqrt{-1 + \frac{\Phi[r]^2}{\Theta[r]^2}} \Omega[r]$  for tensor A

Added independent rule  $\mathcal{A}_2 \rightarrow 0$  for tensor A

Added independent rule  $\mathcal{A}_3 \rightarrow 0$  for tensor A

$$\mathcal{A}_\alpha \quad (9)$$

$$\{\mathcal{A}_0, \mathcal{A}_1, \mathcal{A}_2, \mathcal{A}_3\} \quad (10)$$

$$\left\{ \Phi[r], \sqrt{-1 + \frac{\Phi[r]^2}{\Theta[r]^2}} \Omega[r], 0, 0 \right\} \quad (11)$$

Now check that a simple expression is carried properly to the component form.

$$g^{\alpha\beta} \left( \dot{\nabla}_\alpha \phi \right) \left( \dot{\nabla}_\beta \psi \right) \quad (12)$$

$$-\frac{\phi'[r] \psi'[r]}{\Omega[r]^2} \quad (13)$$

Next something more advanced: we want to show that our one-form really does the job.

$$\mathcal{A}_\alpha \mathcal{A}^\alpha \quad (14)$$

$$\mathcal{A}_\alpha \mathcal{A}_{\alpha'} g^{\alpha\alpha'} \quad (15)$$

$$1 \quad (16)$$

Great. Now let's check that this can be used for more complex expressions.

$$\mathcal{A}_\alpha \mathcal{A}_\beta R[\overset{\circ}{\nabla}]^{\alpha\beta} \quad (17)$$

$$\mathcal{A}_\alpha \mathcal{A}_\beta g^{\alpha\alpha'} g^{\beta\beta'} R[\overset{\circ}{\nabla}]_{\alpha'\beta'} \quad (18)$$

$$\frac{2\Phi[r]^2(\Omega[r]\Theta'[r] + \Theta[r]\Omega'[r]) + \Theta[r]^2(-2\Theta[r]\Omega'[r] - r\Theta'[r]\Omega'[r] + \Omega[r]r\Theta''[r])}{\Theta[r]^3\Omega[r]^3r} \quad (19)$$

### Connection to Part III project

Brilliant, that computation took a fraction of a second. So with these tools all the field equations can be reduced to ODEs in the radius within no more than a couple of hours' tinkering: it is basically just data entry.